

## Milky Way Galaxy

Determining the size/extent – counting stars (doesn't work)

### Variable Stars

Red Giants/Supergiants

Instability Strip – Hydrostatic Equilibrium

Cepheids – characteristics

Type I, II – differences

Leavitt Law (period-luminosity relation) – used to get distances

RR Lyrae - characteristics

### Clusters

Open Clusters - properties

Globular Clusters – properties

Ages of Clusters

### Shapley's Model of Milky Way

Using globular clusters to determine our distance

### Current Model for the Galaxy – major components

Bulge

Disk

Halo

Kiloparsec (kpc) = 1000 parsecs

### Galactic Bulge

Center of the galaxy – direction of Sagittarius

Methods of seeing the center

Object at the Center

Sgr A\*

Massive black hole

### Disk

Characteristics

H I - 21 cm radio signal – main disk material

Spiral Arms

Finding location/number of spiral arms - Spiral Arm tracers

Very sloppy arms

Galactic rotation/motion of the disk

Rotation curve for the disk

Sun's motion around galaxy

Determine mass of galaxy (using Kepler's 3<sup>rd</sup> law)

Evidence for Dark Matter

Spiral Pattern motion

Spiral Density Wave theory (the nude-jogger)

### Halo

Globular Clusters

Dark Matter

Streams of material

### Stellar Populations

Population I Characteristics

Population II Characteristics

## Galaxies

Curtis-Shapley Debate (Harlow Shapley vs Heber Curtis)

Size of the galaxy

Nature of spiral nebula

Edwin Hubble, Milton Humason – distances to galaxies

Standard Candles

Definition

Examples

Best ones

Hubble Classification Scheme for Galaxies

Tuning fork diagram

Ellipticals

Criteria for classification (E0 – E7)

Characteristics

Giant Ellipticals (cD types)

Dwarf Ellipticals and Dwarf Spheroidals

Spirals

Criteria for Classification (Sa, Sb, Sc, Sd, SBa, SBb, SBc, SBd)

Characteristics

Barred

SO (Lenticular)

Characteristics

Barred

Irregular

Galaxy Names and Catalogs

Galaxy Velocity

Hubble Law

$v=H_0d$

Hubble Diagram – velocity vs distance for galaxies

Hubble Constant =  $H_0$

Importance of Hubble's Law

Measurement of distances

Expansion of the Universe

Age of the Universe

Mass determination

Rotation Curve method

Binary motion

Mass-Luminosity Ratio

Galaxy evolution

Angular momentum

Other effects

Galactic Clusters

Local Group

Rich vs Poor clusters

Megaparsec (Mpc) = a million parsecs

Dark Matter

- WIMPs
- MACHOs
- Superclusters
  - Local Supercluster
  - Clusters of Superclusters (?)
- Larger scale structures
  - Voids
  - Concentrations of galaxies
- Active Galaxies
  - Active galactic nuclei
  - Criteria for active galaxies
  - BL Lac
    - Variable brightness
    - Strong central source
    - Non-thermal light
    - Flat spectra
  - Seyfert Galaxies
    - Variable brightness
    - Strong central source
    - Emission lines
    - Type I and Type II
  - Radio Source Galaxies
    - Lobes
    - Jets
    - Cen A, M87
- Quasars
  - High redshift = high velocities = relativistic velocities
  - Emission lines
  - Great distance implied by Hubble's law
  - Variable brightness
  - Small energy source
- Gravitational Lensing
- Unified model for active galaxies
  - Massive Black Holes
  - Accretion Disk
  - High velocity clouds
  - Low velocity clouds
  - Dusty torus
  - Jets
  - Evidence for model
- Most distant objects
  - Galaxy evolution information
  - Hubble Deep Fields
- Cosmology
  - Newton's Universe - Static, Infinite
  - Olber's Paradox - solutions

- Observable Universe
- Properties of the Universe -
  - Homogeneity, Isotropy, = Cosmological Principle
  - Universality
- Uniform expansion of the Universe
  - Curvature of the Universe
    - Flat
    - Positive (spherical)
    - Negative (hyperbolic)
- Einstein's view of the Universe
  - Introduced the Cosmological Constant
- Age of the Universe
  - Methods to estimate age
- Big Bang - Cosmogony
  - Big Bang Theory – Aleksandr Friedmann, George Lemaitre
  - Before the big bang?
    - Time =0
    - Supergravity
  - Planck Era
    - GUT
  - Unified Forces
  - Creation of matter
  - Inflation Era
  - Heavy Particle Era
  - Light Particle Era
  - Nucleosynthesis Era
  - Recombination (decoupling)
  - Galaxy Formation
    - Top-down vs bottom-up
    - Hot dark matter vs cold dark matter
  - Now
- Support for the Big Bang theory
  - Cosmic Background Radiation
    - Penzias and Wilson
    - 3 K
  - Later measurements
    - COBE
    - BOOMERanG
    - WMAP
- Fate of the Universe
  - Gravity
    - Critical Density
  - Options:
    - Open – not dense enough
    - Closed – too dense
    - Flat (marginally open) – exactly equal to critical density

## Current View

Accelerating Universe

Dark Energy, Einstein's Cosmological Constant

Big Rip?

Currently observed characteristics of the Universe

13.7 billion

Flat

Mainly Dark energy + dark matter